AI FOR COSMOLOGY

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INTRODUCTION



This talk highlights a subset of the AI-for-Cosmology efforts at Argonne.

Common themes here are:

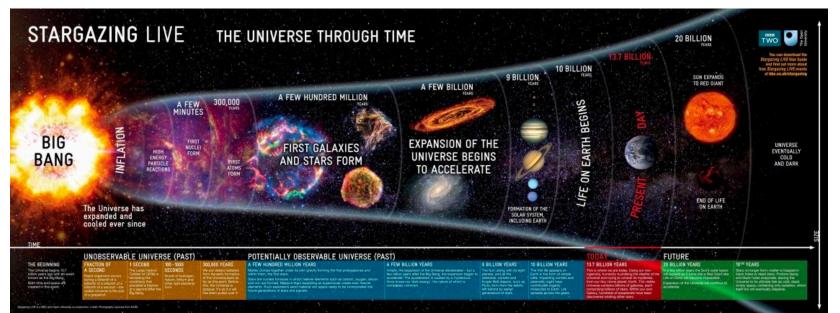
- Synthetic/simulation data to enhance/replace real astronomical observations.
- Bayesian/probabilistic schemes rather than point-predictions.
- Explainability of the Al algorithms.

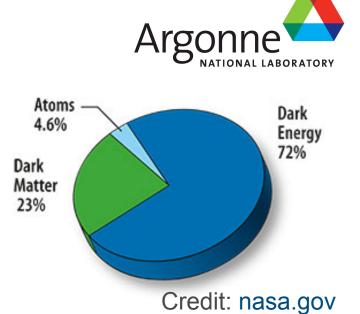
Case study:

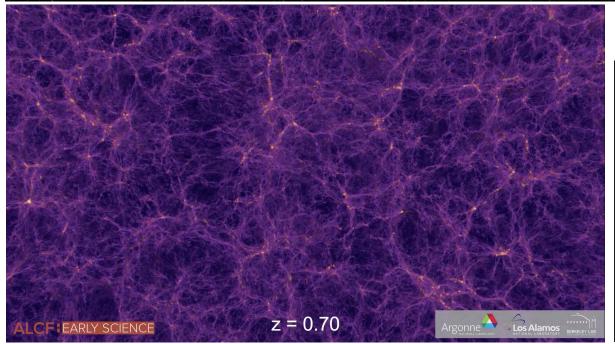
- Image processing pipelines for de-noising, de-blending etc.
- Probabilistic classification and regression
- Latent space exploration



SHORT INTRO TO COSMOLOGY







Source: open.edu

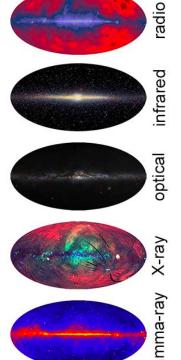
Name (Symbol)	Description	Value
Hubble Constant (H₀)	Current rate of expansion of the universe.	67.4 ± 0.5 km/s/Mpc
Cosmological Constant (A)	Energy density of space, or vacuum energy.	$1.1056 imes 10^{-52}$ m $^{-2}$
Dark Energy Density (ΩΛ)	Fraction of the universe's energy density consisting of dark energy.	0.6847 ± 0.0073
Dark Matter Density (Ωc)	Fraction of the universe's energy density consisting of dark matter.	0.264 ± 0.013
Baryon Density Parameter (Ωb h²)	Density of ordinary matter (baryons) relative to the critical density.	0.0224 ± 0.0001

STUDYING THE UNIVERSE: JOINT EFFORTS



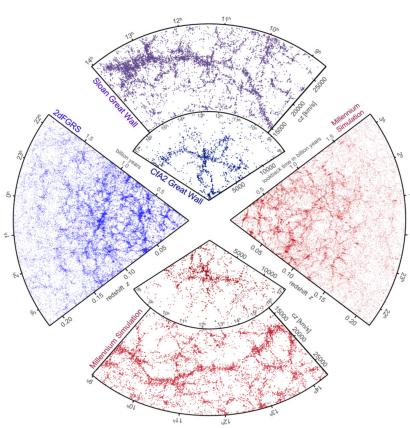




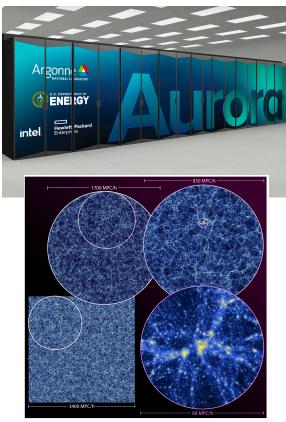




Observational Astronomy



Zavala, J.; Frenk, C.S. Dark Matter Haloes and Subhaloes. *Galaxies* **2019**, *7*, 81.



Theoretical and Computational Cosmology

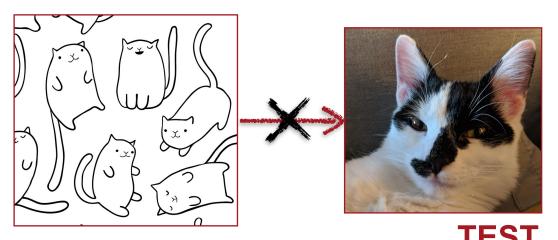


CAVEATS IN COSMOLOGICAL STUDIES



- Multi-modal data:
 - Graphs, images, time-series, summary vectors, scalars, text.
- Expensive:
 - Both simulations and observations are from expensive science campaigns
- Multi-fidelity:
 - Data from different sources have different resolutions, approximations and systematic effects.
 - Transfer of knowledge is not straightforward.
- Data coverage:
 - Gaps, biased datasets are common.
 No assumption of a 'fair' sampling.

- Analysis requirements
 - Precision cosmology has high error requirements
 - Traditional statistical analyses have been highly successful.
- Prior domain knowledge:
 - Studies assume known physics, conservation laws.

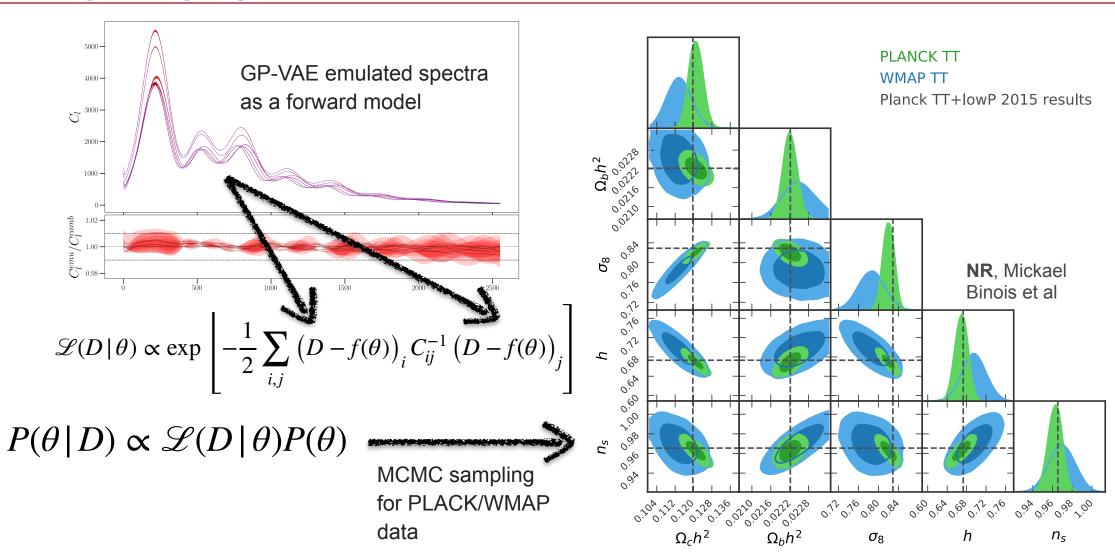






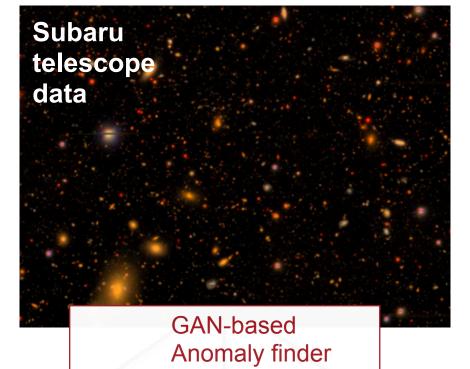
A FEW EXAMPLES: BAYESIAN INFERENCE WITH EMULATORS



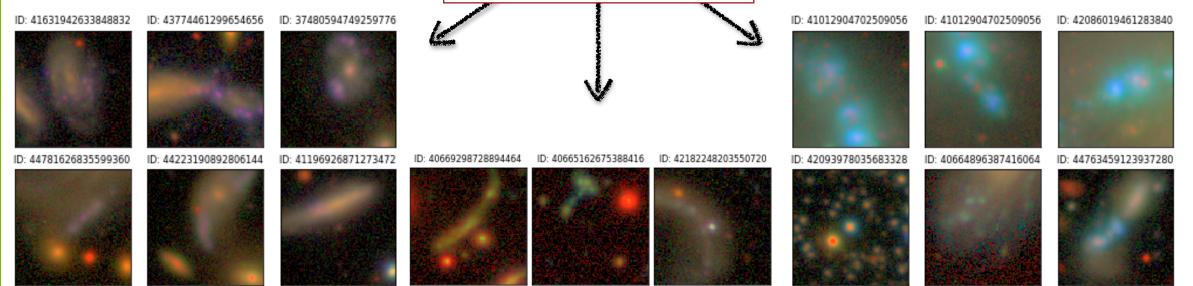


A FEW EXAMPLES: FINDING UNKNOWN UNKNOWNS OBJECTS

IN THE SKY

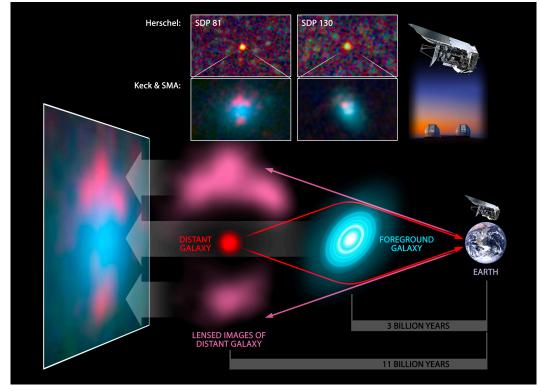


https://arxiv.org/abs/2012.08082



CASE STUDY: GALAXY-SCALE STRONG LENSING

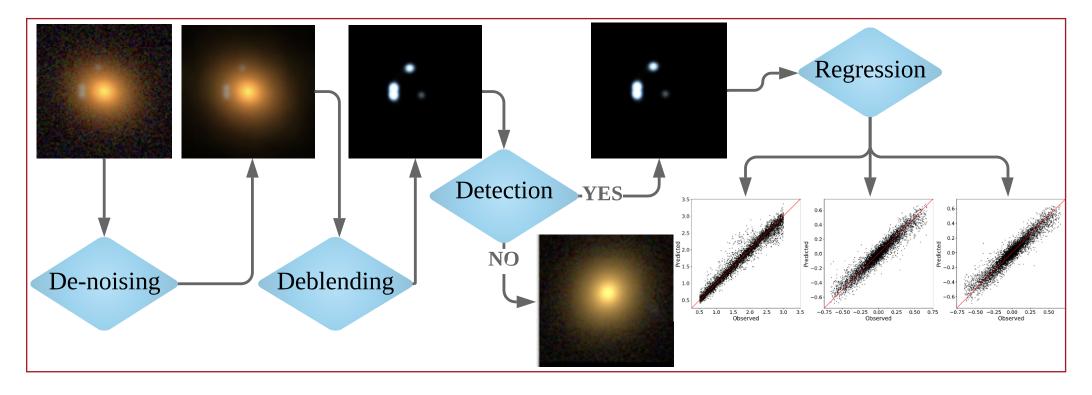
- Strong-lenses are rare objects.
 - But understanding them is key to several questions: Distribution of dark matter, expansion of the Universe.
- Discrepancy with current amount of observed data vs future data
 - Observed data is/will be a highly imbalanced dataset
 - Tractable physical models





INTERPRETABLE LEARNING PIPELINES





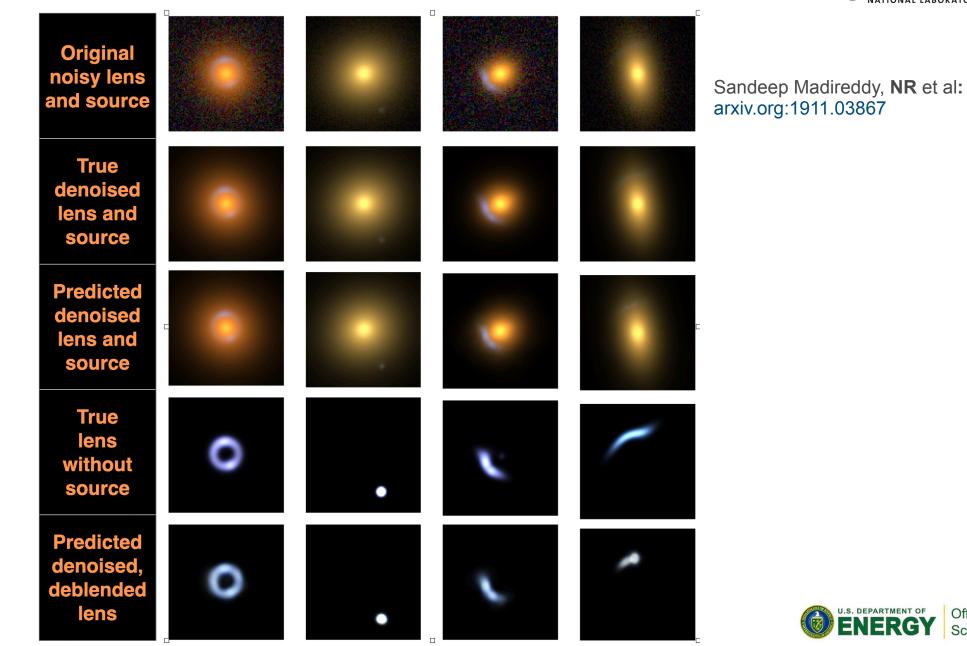
- Each Al-module is independently trained and validated. (Super-resolution modules for Denoising and deblending, Information bottleneck design for detection and regression)
- Synthetic data allows one to train modular pipelines that enable better control over systematics than end-to-end training methods



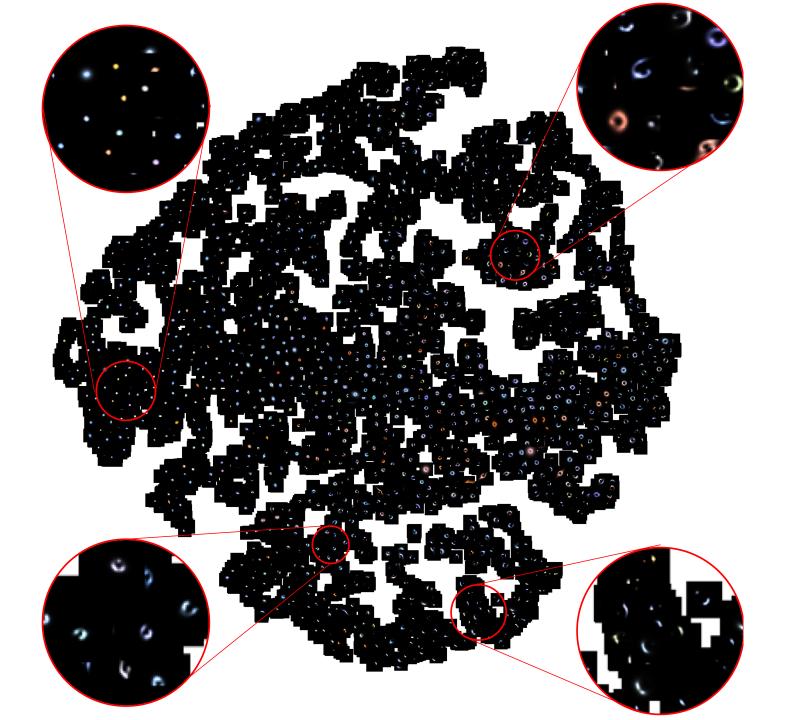
INTERPRETABLE STRONG LENS END-TO-END ANALYSIS Argonne

PIPELINE



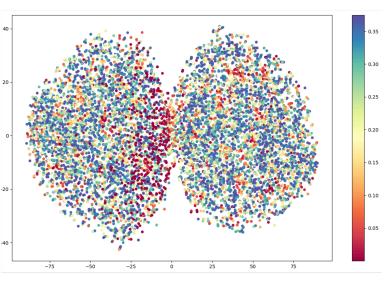








Variational
Information
Bottleneck and
representation
learning



Uncertainty quantification for classification



CONCLUSIONS



- Cosmological studies involve variety of data modalities, with vast amount of data. This makes data-driven Al-models extremely valuable.
- Synthetic datasets are often a necessity in Cosmological analysis.
- Careful experimental design, robust data creation, extensive validations are all required while dealing with synthetic data.
- Interpretable, uncertainty quantified models are still very important, probably even more so while using synthetic data in training.

